Current status of the Fly's Eye Camera System

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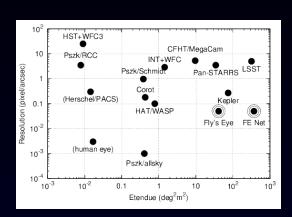
The Fly's Eye Camera System

- 19 wide field cameras with Sloan filters with 22"/pixel resolution, 26° FoV per unit
- Hexapod mount for sidereal tracking
- Autonomous operation, weatherproof enclosure



Scientific Goals

- Time-domain astronomy
- All-sky survey with high cadence and étendue (FoV multiplied by effective light collecting area[deg²m²])
- Planetary system development
- Star formation and evolution
- Extragalactic phenomena



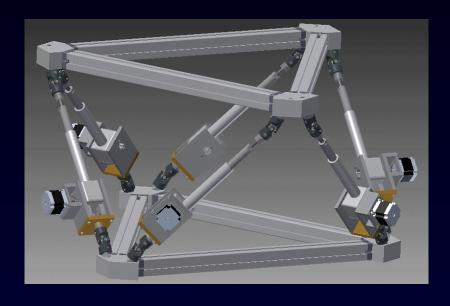


Hexapod

- 6 degrees of freedom,
 only 3 for tracking
- →3 legs would stuck,
 tracking is still manageable
- Tracking drift: ~0.5"min⁻¹
- Self-calibration: independent from longitude and latitude (no polar alignment required)
- Also no leveling required

Hexapod

- Linear actuators (0.05µm stroke per motor step)
- Redundant monitoring





Cameras & Lenses

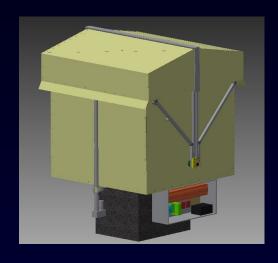
- 19 wide-field cameras with Sloan filters ('g/'r/'i optional 'u/'z)
- Lenses: f=85mm, f/1.2
- Photometric precision:~4-5mmag
- Limit: ~9^m r ≤ ~15^m (close to LSST saturation limit)





Enclosure

- Custom-designed waterproof enclosure
- Doors are moved by outdoor linear actuators

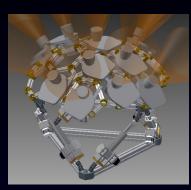


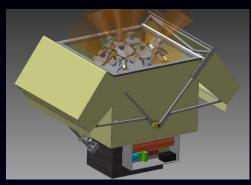


Current Status

- Hexapod motion control ✓
- Camera rack ✓
- Camera units ✓
- Enclosure ✓
- Software ✓







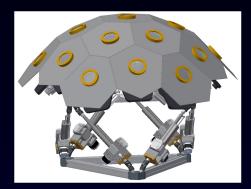


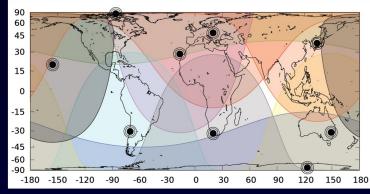
Future plans

- Isolated enclosure, adjustable temperature, humidity within
- Fly's Eye Net: 8-9 units could monitor the whole sky in r<15^m regime (total

étendue ~LSST)







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Current status of the Fly's Eye Camera System

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This project aims to provide a low resolution and multiple-passband full-sky survey with an imaging cadence of a few minutes. Based on our earlier tests, we found that a nove type of astronomical telescope mount on a hexagod platform can provide the accuracy of sidereal tracking needed by our instrumentation. The fully configured Fly's Eye device contains 19 wide-field cameras equipped with fast focal ratio optics which are arranged scientific goal of this project is to continuously monitor stellar brightness variations in the full Sloan photometric system down to the magnitude of r=15 or the limit of apparen stellar confusion. The data acquisition will then cover roughly 6 magnitudes of the time domain, from the scales of minutes up to the several years of planned operations. Fly's Eye data yield is complementary to that of the Large Synoptic Survey Telescope since the saturation magnitude of LSST is close to the faint limit of the Fly's Eye setup.

One of the main scientific yields of this survey is to recover time-domains of photometric variability of stars with magnetic activity. These timescales range from minutes through hours to years, just like in the case of the Sun. If active stars are monitored continuously the measurements will give us data in the broad time range of the magnetic phenomena observed and their spatial correlation studied. The Flv's Eve device allows simila research on different kinds of active stars individually, observing in five band passes from ultraviolet to near-infrared (365 to 900 nm) which is unprecedented. The results give us broader view of the magnetic activity of stars of different ages. Through this, we will be able to reconstruct the past of the Sun and foresee its future



Figure 3. Tracking, using an f=800mm lens during 3min interval (top and lower left). Image stamp of 64x64 pixels, taken with an f=85mm ens, exposure time: 130seconds (lower centre). PSF of the stellar profile at the center of the



Figure 4. An opcional 2nd layer of isolated housing, within it the temperature, humidity can be adjusted. An example configuration of









Figure 1. The Fly's Eve Camera System, 19 camera units are supported by a frame mounted on a hexapod mechanics. Unner-left: CAD model of the system (the brown cones are the aboratory (several cameras are already available) Lower-left: image. Lower-right: Étendu and effective resolution fo









Figure 2. The custom-designed enclosure for the Fly's Eye. Th housing is mounted on a concrete basement. The doors are moved by weatheproof linear actuators. Left: CAD models of the enclosure. Right: The system installed on the concrete basement at Piszkés-tető Observatory, Hungary.

- SCIENTIFIC GOALS
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